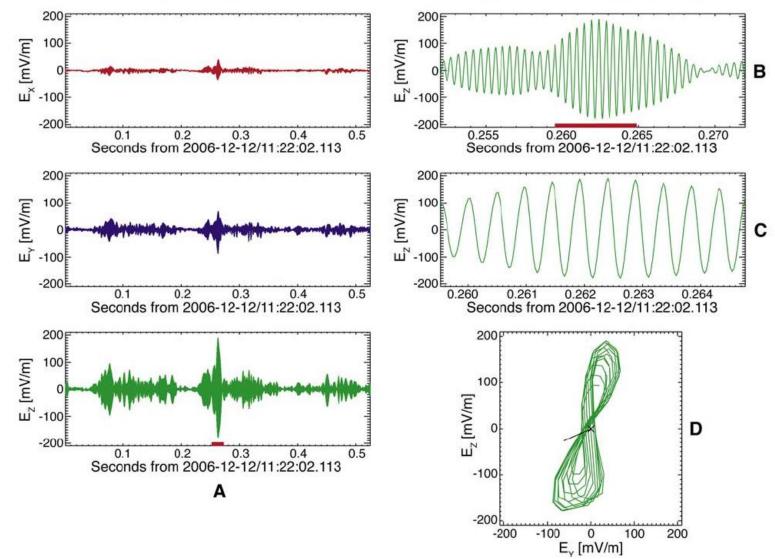
# Large-Amplitude Whistler Waves and Relativistic Electron Acceleration

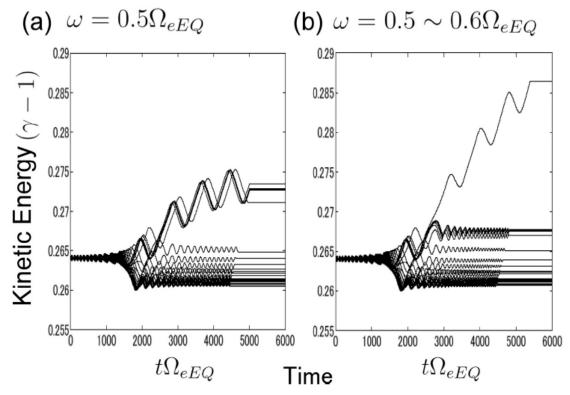
Peter H. Yoon (Kyung HeeUniv&Univ of Maryland)





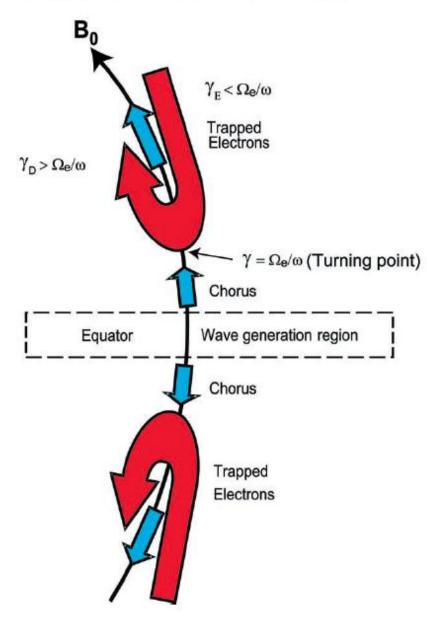
Cattell et al. (2008): Large-amplitude (200 mV/m or  $\delta B/B_0 \approx 0.01$ ) oblique (40°-70°) whistler wave observed at high magnetic latitude (~ 20°).

## **Theories of Whistler-Acceleration**



- Omura and Summers (2006):
  - Inhomogeneous B field. Resonant trapped electrons forming electromagnetic electron hole. Coherent whistler mode. Acceleration of trapped electrons.

#### RELATIVISTIC TURNING ACCELERATION



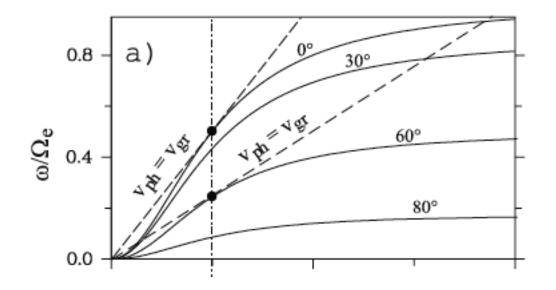
• Omura et al. (2007):

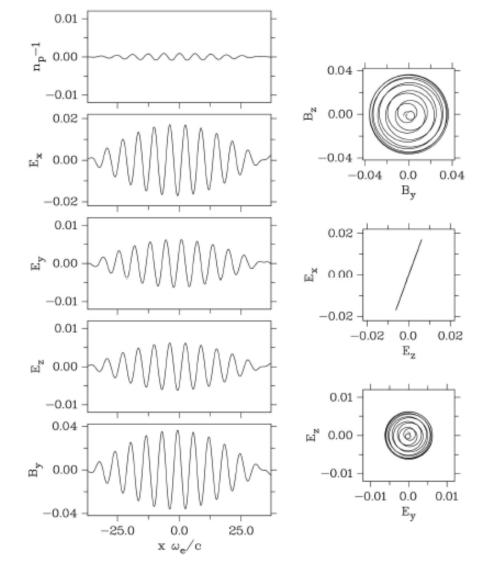
Extension of
 Omura& Summers
 (2006). Relativistic
 turning acceleration

- Omura& Summers (2006) and Omura et al. (2007) assume parallel propagation of whistler wave.
- Cattell et al. (2008) STEREO observation is for oblique propagation angle.

## **Whistler Oscilliton Theory**

- Sauer et al. (2002)
- Dubinin et al. (2003, 2007)
- Sauer and Sydora (2010)
- Coherent wave packets for whistler waves satisfying Gendrin condition.





Sauer and Sydora (2010)

Note: The authors do not discuss electron acceleration.

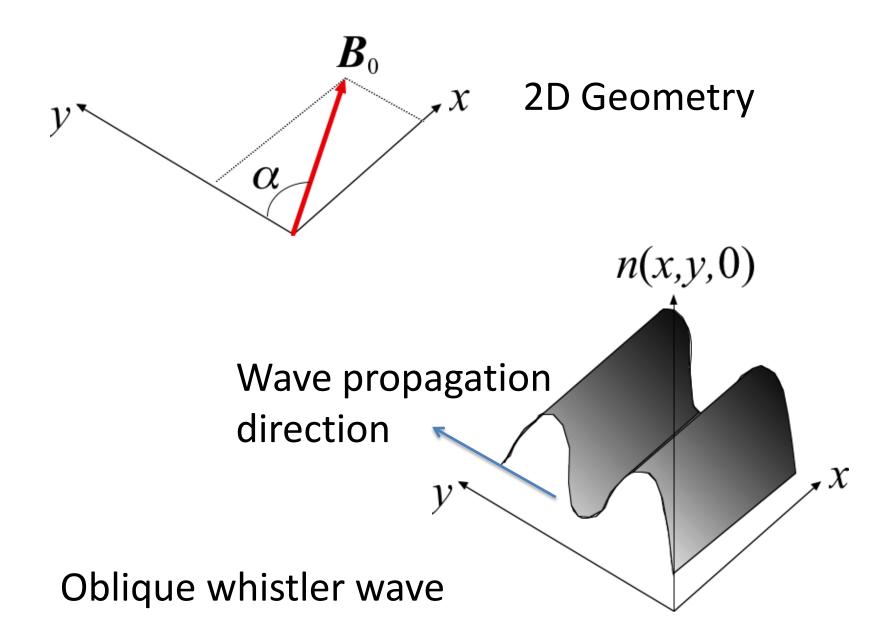
**Fig. 8.** Spatial profiles of whistler oscillitons for  $\theta = 70^{\circ}$  and  $U = 0.172V_{Ae}$ . From top to bottom: proton density  $(n_p - n_{po})/n_{po}$ , three components of the electric field (in units of  $E_o = V_{Ae}B_o$ ) and the magnetic field component  $B_y/B_o$ . The right panels show the hodographs  $B_z$  versus  $B_y$ ,  $E_x$  versus  $E_z$ , and  $E_z$  versus  $E_y$ .

# **Oblique whistlers**

 Tao and Bortnik (2010): Theory of oblique whistler and relativistic electron acceleration — using linear theory of oblique whistler wave.

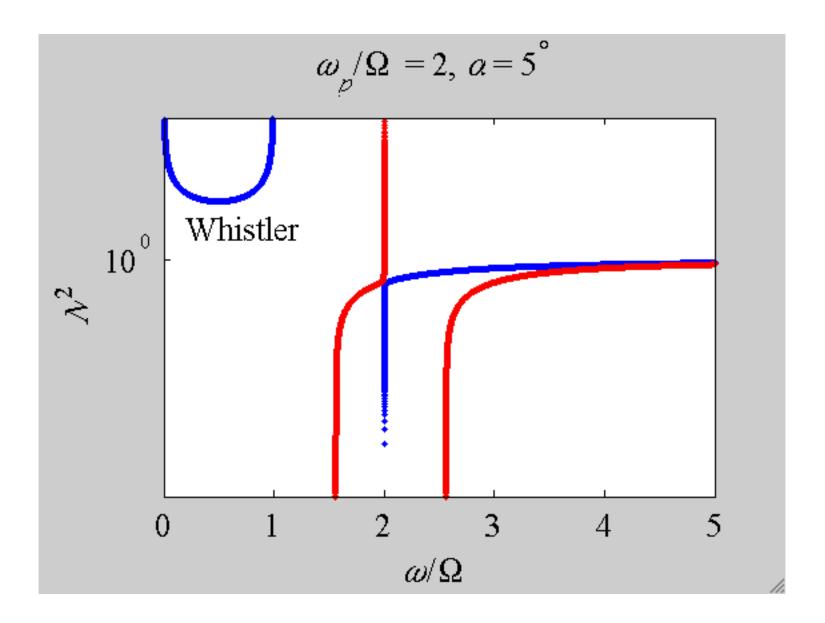
 Verkhoglyadova et al. (2010): Linear theory of oblique whistler wave — they do not discuss relativistic electron acceleration.

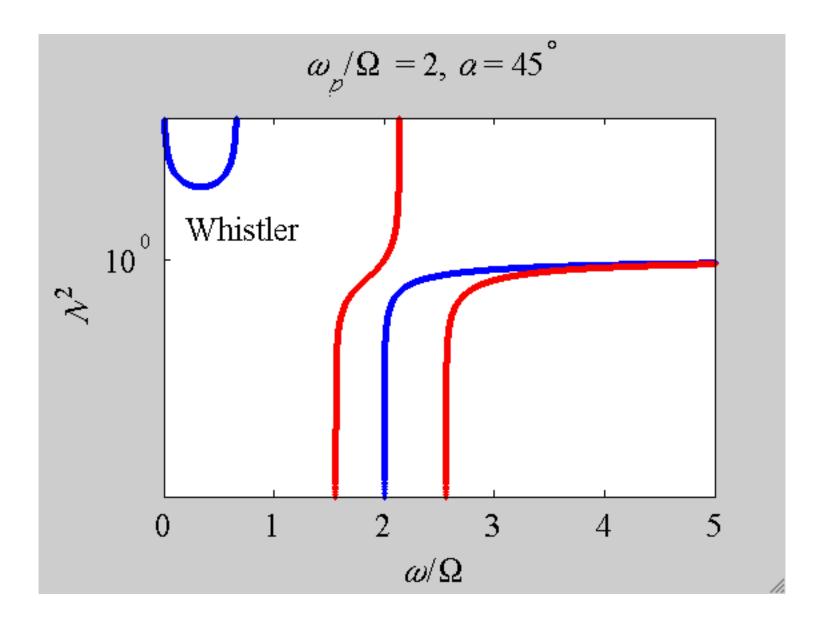
## Nonlinear oblique whistlers

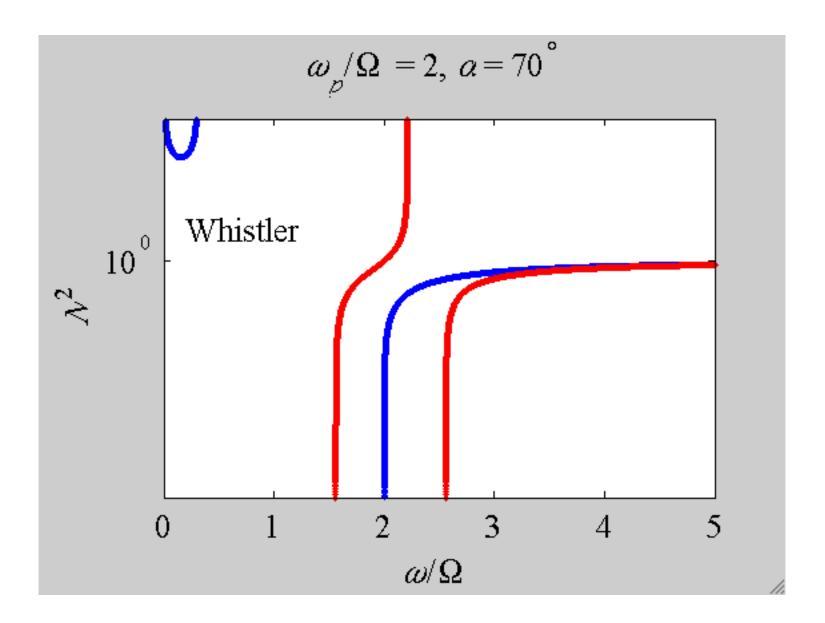


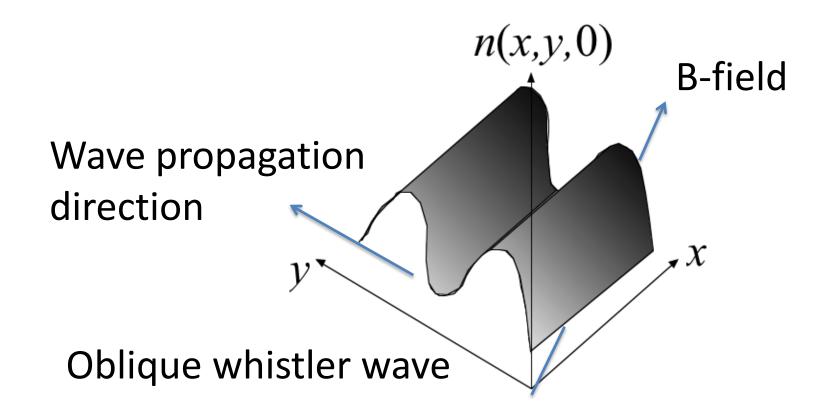
## Initial condition

## **Oblq Whist Wave Dispersion Relation**

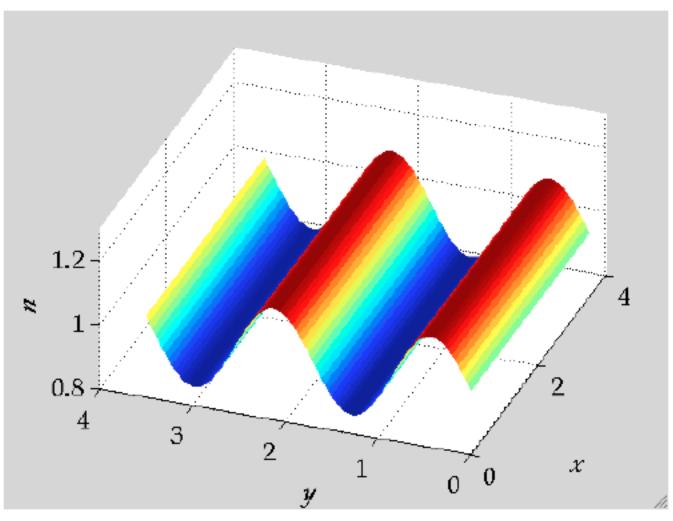


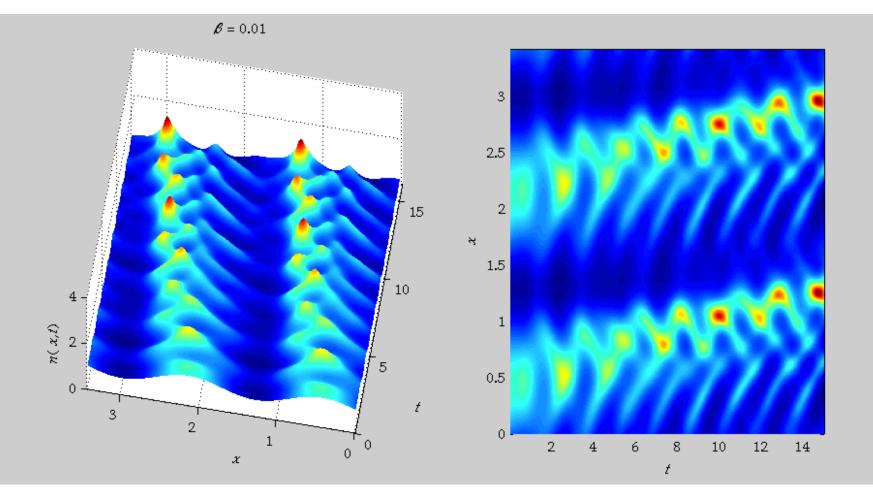




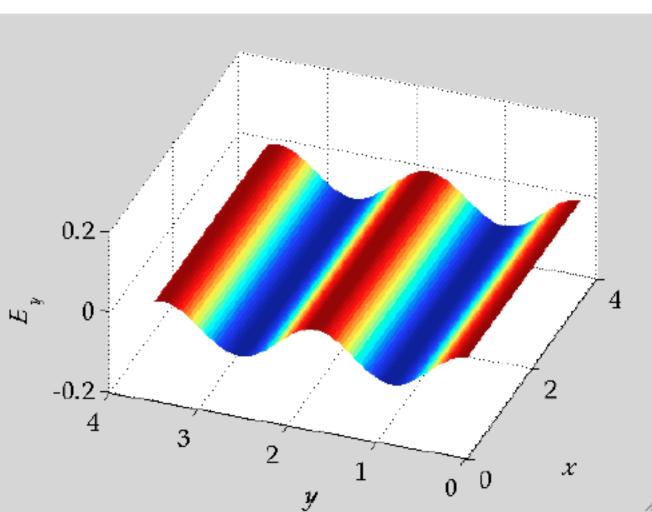


T=0.1

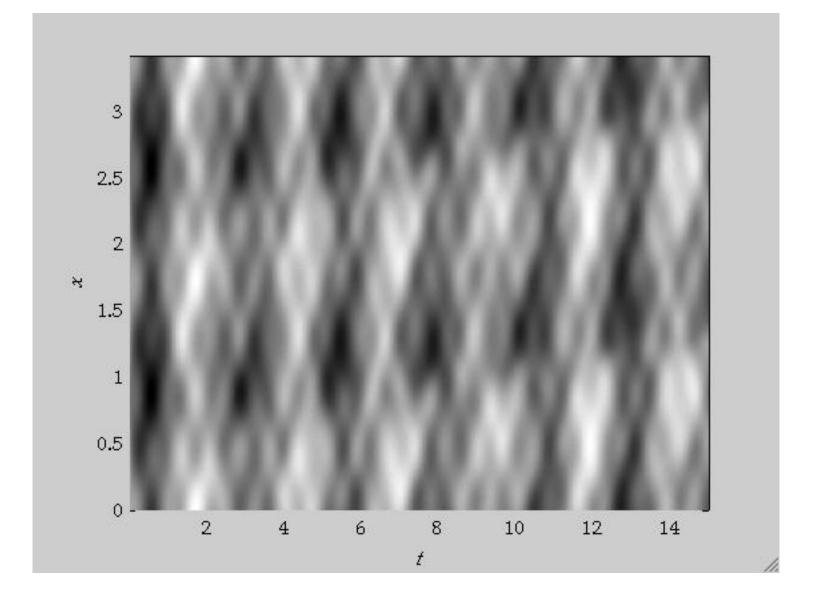




11.



T = 0



## **Relativistic e-acceleration**

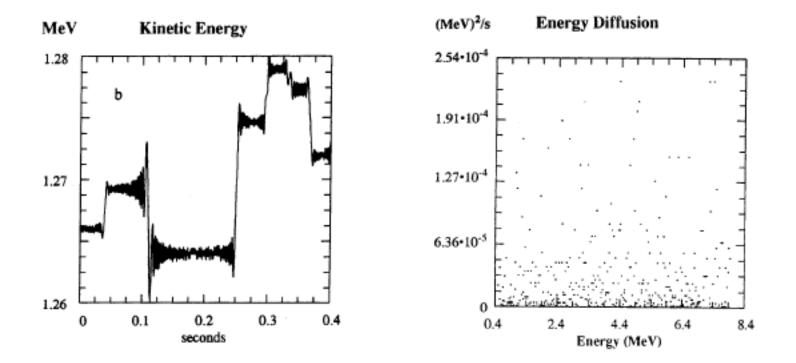
Test-particle simulation over nonlinear whistler wave (cf., Roth et al., 1999)

### Resonant enhancement of relativistic electron fluxes during geomagnetically active periods

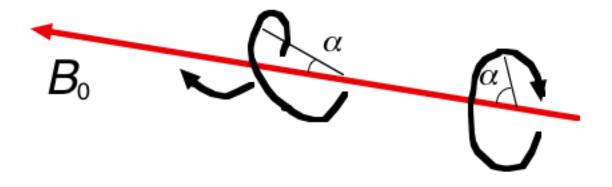
#### I. Roth<sup>1</sup>, M. Temerin<sup>1</sup>, M. K. Hudson<sup>2</sup>

<sup>1</sup> Space Sciences Laboratory, University of California, Berkeley, CA 94720, USA e-mail: ilan@ssl.berkeley.edu
<sup>2</sup> Physics and Astronomy Department, Dartmouth College, Hanover, NH 03755, USA

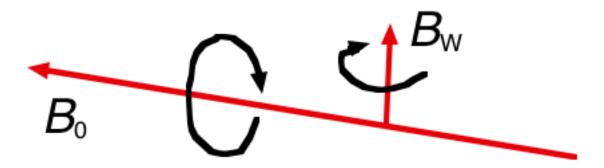
Received: 30 June 1998 / Revised: 26 October 1998 / Accepted: 27 October 1998

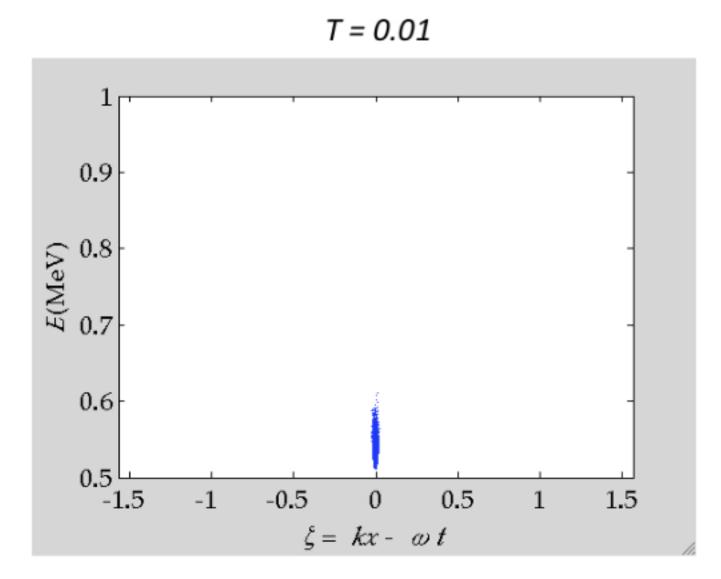


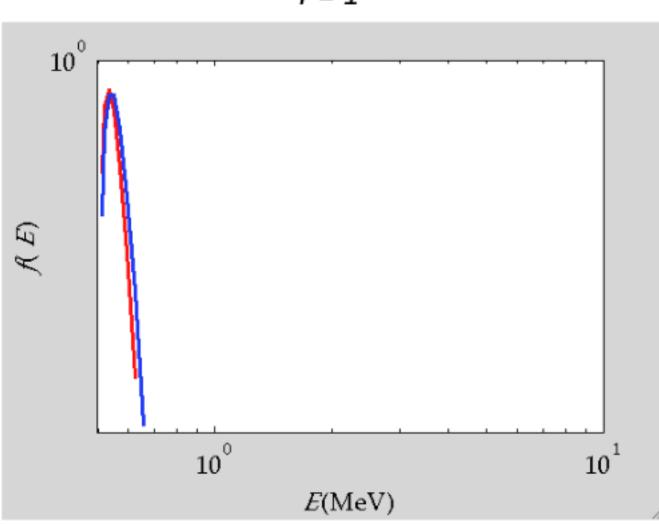
### Wave-particle resonance



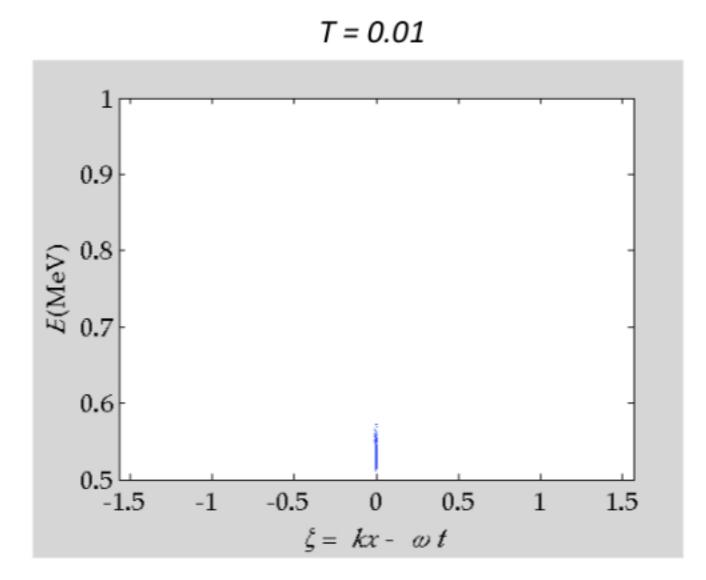
Wave-particle scattering (includes non-resonant wave-particle interaction)

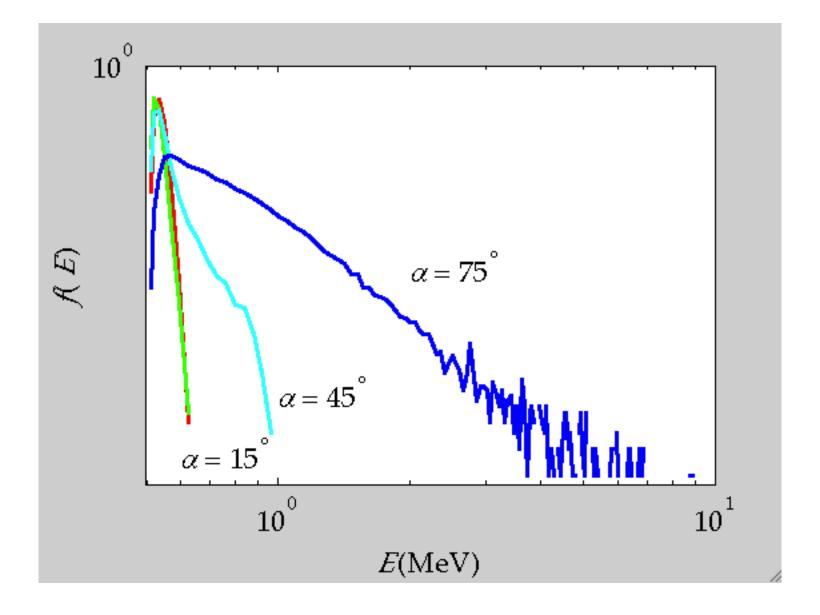






T = 1





# **Conclusion and Discussions**

- Rapid acceleration of relativistic electrons by obliquely-propagating whistler waves.
- The problem of how these waves are created is outstanding.
- Assuming that the source of oblique whistlers are near the equator, the propagation is not understood.